

# Firehose and Mirror instabilities

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Cold electron CGL MHD linear dispersion relation

$$\omega^4 - \omega^2 \underbrace{\left[ \left( \frac{2P_{\perp}}{\rho_0} + V_A^2 \right) k_{\perp}^2 + \left( \frac{2P_{\parallel}}{\rho_0} + \frac{P_{\perp}}{\rho_0} + V_A^2 \right) k_{\parallel}^2 \right]}_{A_2} + 3k_{\parallel}^2 \frac{P_{\parallel}}{\rho_0} \underbrace{\left[ \left( \frac{2P_{\perp}}{\rho_0} - \frac{P_{\perp}^2}{3P_{\parallel}\rho_0} + V_A^2 \right) k_{\perp}^2 - \left( \frac{P_{\parallel}}{\rho_0} - \frac{P_{\perp}}{\rho_0} - V_A^2 \right) k_{\parallel}^2 \right]}_{A_0} = 0$$

$$V_A^2 = \frac{\beta^2}{\mu_0 \rho_0} \quad \rho_0 \text{ density} \quad P_{\perp}, P_{\parallel} \text{ pressures}$$

$$k_{\parallel} = \hat{b} \cdot \underline{k} \quad \underline{k}_{\perp} = \underline{k} - k_{\parallel} \hat{b}$$

Parallel  $k_{\perp} \ll k_{\parallel}$

$$\omega^2 = \frac{1}{\rho_0} \left( \frac{\beta^2}{\mu_0} - P_{\parallel} + P_{\perp} \right) k_{\parallel}^2 \quad \text{Alfvén wave}$$

if  $P_{\perp} = P_{\parallel} \Rightarrow$  MHD shear Alfvén wave

$$\text{if } P_{\parallel} > P_{\perp} + \frac{\beta^2}{\mu_0} \Rightarrow \omega^2 < 0$$

instabilities

$$\beta_{\parallel} = P_{\parallel} / (\beta^2 / \mu_0) \quad \beta_{\perp} = P_{\perp} / (\beta^2 / \mu_0)$$

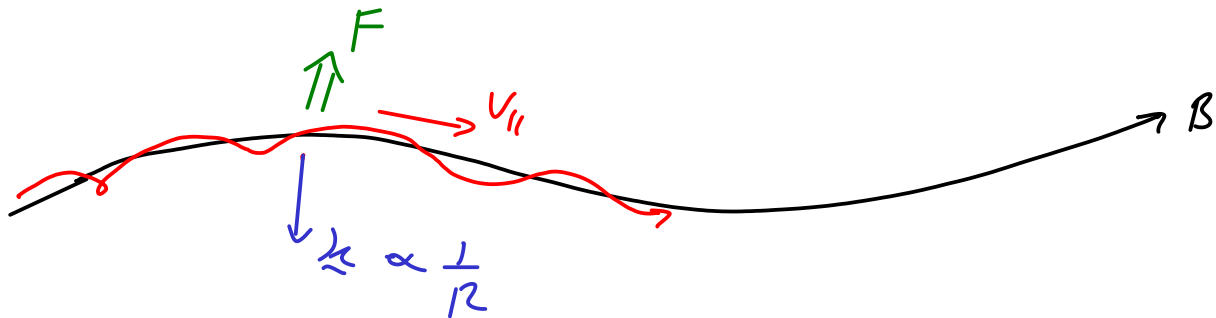
$$\underline{\underline{\beta_{\parallel} > \beta_{\perp} + 2}}$$

$$\underline{\underline{\beta_{\parallel} > 2}} \quad (\text{if } \beta_{\perp} = 0)$$

$\Rightarrow$  space / astrophysics

not in magnetic confinement fusion

Mechanism: "Centrifugal" instabilities



Centrifugal force  $\rightarrow$  particle drift

$\rightarrow$  can enhance perturbation (instabilities)

Stabilised by magnetic tension or perpendicular pressure

Mirror instabilities

Perpendicular "slow" wave  $k_{\perp} \gg k_{\parallel}$

$$\omega^4 - A_2 \omega^2 + A_0 = 0$$

$$\omega^2 = \frac{1}{2} \left( A_2 \pm \underbrace{\sqrt{A_2^2 - 4A_0}}_{>0} \right) \quad A_2 > 0$$

if  $\sqrt{A_2^2 - 4A_0} > A_2 \Rightarrow$  instability

$\Rightarrow A_0 < 0$  unstable

$$k_{||}^2 \left( 1 - \frac{1}{2} \beta_{||} + \frac{1}{2} a \beta_{||} \right) + k_{\perp}^2 \left( 1 + a \beta_{||} - \frac{1}{6} a^2 \beta_{||} \right) < 0$$

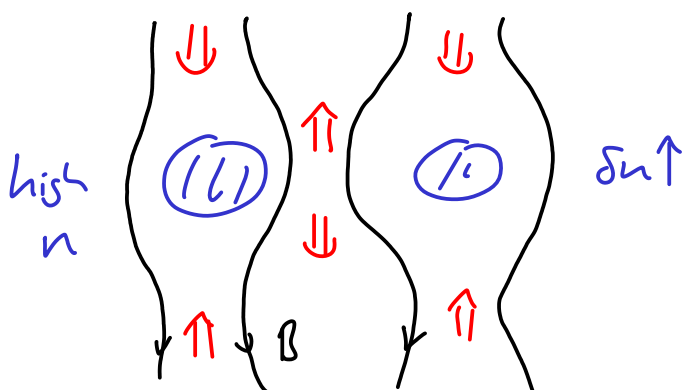
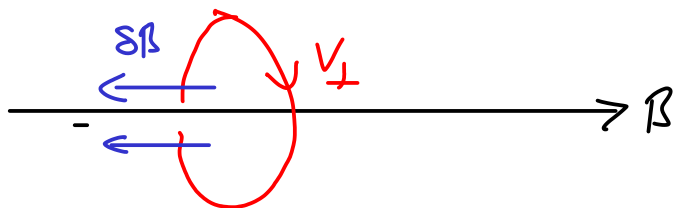
$a = \frac{T_{\perp}}{T_{||}}$  anisotropy

$\frac{\beta_{\perp}^2}{\beta_{||}} > 6(1 + \beta_{\perp})$

instability

$= 1$  in kinetic theory

Mechanism: "diamagnetic" instabilities



Perpendicular pressure  
reduces B field  
 $\Rightarrow$  creates a potential  
well and more  
trapping

$\delta B, \delta n$  anti-correlated

