## Small-scale instabilities

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## This lecture

- We've looked at MHD instabilities, which tend to be the fastest and most dangerous for confinement
- In this lecture we'll look at other instabilities which degrade confinement but don't lead to catastrophic results.
- These result from treating electrons and ions separately and so are two-fluid or kinetic effects
- They are essentially electrostatic, and are driven by gradients in temperature and density
- These are thought to be the origin of turbulence in confinement devices, and so anomalous transport

References:

- J. Wesson "Tokamaks", sections 8.2 8.5
- J.W.Connor, H.R.Wilson "Survey of theories of anomalous transport" *Plasma Phys. Control. Fusion* **36** 719-795 (1994)
- B.D.Scott "Computation of turbulence in magnetically confined plasmas" *Plasma Phys. Control. Fusion* 48 B277 (2006)

- Because electrons move quickly along magnetic fields, they are often assumed to quickly reach equilibrium on the timescale of instabilities.
- This simplifies the analysis as we can concentrate on the ions, and assume that the electrons follow.

The momentum equation for electrons is

$$n_e m_e \left( \frac{\partial \underline{v}_e}{\partial t} + \underline{v}_e \cdot \nabla \underline{v}_e \right) = -\nabla p_e - n_e e \left( \underline{E} + \underline{v}_e \times \underline{B} \right)$$

Parallel to the magnetic field, and setting the left side to zero

$$n_e e E_{||} + \nabla_{||} p_e = 0$$

#### Electron response

Linearising this equation, neglecting temperature variations because parallel thermal conduction is fast (so  $\nabla_{||} T \simeq 0$ ), and assuming electrostatic perturbations so  $E_{||} = -\nabla_{||}\phi$ 

$$-(n_0+\delta n) e\nabla_{||} (\phi_0+\delta \phi) + \nabla_{||} [T_0 (n_0+\delta n)] = 0$$

$$-n_0e\nabla_{||}\delta\phi - \delta ne\nabla_{||}\phi_0 + \nabla_{||}(\delta nT_0) = 0$$

Since there are no parallel gradients of equilibrium  $\phi_0$  and  $n_0$ , this becomes

$$-\nabla_{||}(n_0e\delta\phi)+\nabla_{||}(\delta nT_0)=0$$

and so:

$$n_0 e \delta \phi = \delta n T_0 \qquad \Rightarrow \frac{\delta n}{n_0} = \frac{e \delta \phi}{T_0}$$

This is called the adiabatic or Boltzmann response

Consider a slab of plasma with the  $\underline{B}$  into the page, and density increasing from right to left



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• A small density perturbation

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- A small density perturbation
- The electrons move along the field and establish force balance and so  $\delta \phi \simeq \frac{T_0}{e} \frac{\delta n}{n_0}$

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- A small density perturbation
- The electrons move along the field and establish force balance and so  $\delta \phi \simeq \frac{T_0}{e} \frac{\delta n}{n_0}$
- This gives an <u>E</u> × <u>B</u> velocity which is 90° out of phase with the density
- This is a wave which propagates perpendicular to ∇n and <u>B</u>
- Because δ<u>ν</u> and δn are out of phase, there is no net radial transport.

We can derive the dispersion relation of this wave using the ion density (continuity) equation

$$\frac{\partial \delta n_i}{\partial t} = -\nabla \cdot \left[ \left( n_0 + \delta n_i \right) \delta \underline{v} \right]$$

As before, put in a solution  $\delta n_i \propto \exp(-i\omega t)$ . so that  $\frac{\partial \delta n_i}{\partial t} \rightarrow -i\omega$ . For the ions, assume that radial  $\underline{E} \times \underline{B}$  is the dominant motion (not true for electrons!)

$$\Rightarrow -i\omega\delta n_i = v_{E\times B,r}\frac{dn_0}{dr}$$

The radial  $\underline{E} \times \underline{B}$  velocity is given by the poloidal gradient of the electrostatic potential (since *B* field is mainly toroidal). Taking a single wave of the form exp $(ik_{\theta}r\theta)$ 

$$v_{E \times B, r} = -\frac{1}{B} \frac{\partial \delta \phi}{\partial r \theta} = -\frac{1}{B} i k_{\theta} \delta \phi \Rightarrow n_{i} = \frac{k_{\theta} \delta \phi}{\omega B} \frac{dn_{0}}{dr}$$

Using quasi-neutrality,  $n_i \simeq n_e$  and so

$$\frac{k_{\theta}\delta\phi}{\omega B}\frac{dn_{0}}{dr}=\frac{n_{0}e\delta\phi}{T_{0}}$$

The wave frequency is therefore

$$\omega = \frac{k_{\theta}T_0}{eBn_0}\frac{dn_0}{dr} = \frac{k_{\theta}}{eBn_0}\frac{dp_0}{dr} = k_{\theta}v_* = \omega_*$$

This velocity  $v_*$  is the **diamagnetic drift** velocity, and is the reason why this category of waves are known as **drift waves**.

### Electron drift wave growth

- If velocity and density perturbations are out of phase then there is no net radial transport and the wave doesn't grow.
- If the electrons can't keep up with the wave then this leads to a phase shift and growth of the mode
- Finite electron mass, or just about any form of dissipation e.g. resistivity or Landau damping will have this effect.
- This is often called the **Universal instability** because all useful plasmas have density gradients and some dissipation.
- In fact magnetic shear (change in *q* with radius) can stabilise these modes
- In toroidal geometry these waves are destabilised by trapped particles and become unstable above a threshold given by the magnetic shear

Ion Temperature Gradient (ITG or  $\eta_i$ ) mode

- Start with a small fluctuation in temperature
- The sum of curvature and grad-B drifts is

$$\underline{v}_{\nabla B} + \underline{v}_{R} \simeq \frac{\left(v_{\perp}^{2}/2 + v_{||}^{2}\right)}{R\Omega} \underline{e}_{z}$$

so hotter particles drift faster



Figure: D.Applegate's thesis

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- This leads to regions of higher and lower density
- Using the electron adiabatic response, this gives an <u>E × B</u> drift which amplifies the original perturbation





Ion Temperature Gradient (ITG or  $\eta_i$ ) mode

- The ITG mode is stabilised by magnetic shear, but less so than the electron drift wave
- A proper analysis gives a threshold for stability (see figure)
- The mode gets its name because there is a threshold in ion temperature gradient at a given density gradient
- This is thought to be why tokamak profiles are "stiff" - the gradient tend to be fixed.



## Electron Temperature Gradient (ETG)

- Because the ITG instability depends on the ion ∇B and curvature drift, it has wavelengths of a few ion Larmor radii.
- There is another instability which has a scale between the electron and ion Larmor radii.
- Because it is smaller than the ion Larmor radius, the instability doesn't "know" that ions are on orbits. Instead, it sees a Boltzmann-like response for the ions, similar to the electron behavior in the ITG mode.
- The ETG mode is therefore very similar to the ITG mode, but with the roles of the electrons and ions reversed
- Sidenote: there has been a long and ongoing debate over whether ETG or ITG is more important for tokamak transport

## Trapped particle modes

Particle trapping in toroidal plasmas means there are two populations of particles:

- passing particles which have a net parallel velocity and which explore all parts of the torus
- trapped particles which have little net parallel motion, and which only explore part of the torus

This leads to an instability similar to sausage / interchange modes



- For passing particles the good and bad curvature averages out and is stable
- Trapped particles only see the bad curvature side so have a net drift
- The passing particles act like a background with a Boltzmann response

**Problem:** The full 6-D Vlasov equation is too difficult to solve in most situations of interest, but the plasma core is not collisional enough for a fluid (MHD-like) model to be valid

How is plasma turbulence calculated?

- **Gyrokinetics**: Remove fast timescales and reduce number of dimensions
- Numerical tricks: Speed up calculations by many orders of magnitude
- **High Performance Computing**: Algorithms needed to parallelise efficiently across thousands of processors

## Gyrokinetics

- Recall that the Vlasov equation describes a collection of particles, each with a position x and velocity v. Both are 3D, so this is a 6D problem.
- In a strong field, these particles are gyrating quickly ( $\sim$  GHz) around the magnetic field, much faster than the turbulence we want to calculate ( $\sim$  100 kHz)
- We can think of these particles as small current loops



Gyrokinetics describes the dynamics of these current loops

One of the major achievements in plasma theory

- Early work on linear theory e.g. J.B. Taylor, R.J. Hastie (1968), Rutherford and Frieman (1968), P.Catto (1978)
- Nonlinear theory: Frieman & Chen (1982), and first simulations: W. W. Lee (1983)
- Hamiltonian formulation: R.G.Littlejohn (1979,1982), Dublin et al. (1983) ensures conservation of energy
- Modern gyrokinetics uses sophisticated mathematics of differential geometry and field theories

Here I will give only a brief outline of the basic versions

Consider a particle at position  ${\boldsymbol x}$  with velocity  ${\boldsymbol v}$ 

- We need 6 numbers to describe the position of this particle in phase space
- We're free to choose what coordinates to use:

 $(\mathbf{x}, \mathbf{v}) \rightarrow (\overline{\mathbf{x}}, \mathbf{v}_{||}, \mathbf{v}_{\perp}, \phi)$ 

where  $\overline{\mathbf{x}}$  is the middle of the orbit,  $v_{||}$  is the velocity along the magnetic field,  $v_{\perp}$  the speed around the magnetic field, and  $\phi$  is the gyro-phase.

#### Average around an orbit

Averaging over gyro-angle  $\phi$  (gyro-averaging) removes the dependence on  $\phi$ , and reduces the number of dimensions to 5.

• Starting with a distribution of particles *f*, so that the number of particles within a small volume of phase space is

$$\delta n = f(\mathbf{x}, \mathbf{v}) \, \delta \mathbf{x} \delta \mathbf{v}$$

ullet We re-write this in terms of gyro-centre  $\overline{\mathbf{x}}$  and gyro-angle  $\phi$ 

$$\delta \mathbf{n} = \hat{f} \left( \overline{\mathbf{x}}, \mathbf{v}_{||}, \mathbf{v}_{\perp}, \phi \right) \delta \overline{\mathbf{x}} \delta \mathbf{v}_{||} \delta \mathbf{v}_{\perp} \delta \phi$$

• Integrate over gyro-phase

$$\overline{f}\left(\overline{\mathbf{x}}, \mathbf{v}_{||}, \mathbf{v}_{\perp}\right) = \frac{1}{2\pi} \oint \widehat{f}\left(\overline{\mathbf{x}}, \mathbf{v}_{||}, \mathbf{v}_{\perp}, \phi\right) d\phi$$
$$\Rightarrow \delta \mathbf{n} = \overline{f}\left(\overline{\mathbf{x}}, \mathbf{v}_{||}, \mathbf{v}_{\perp}\right) \delta \overline{\mathbf{x}} \delta \mathbf{v}_{||} \delta \mathbf{v}_{\perp}$$

# Equation for $\overline{f}$

To derive an equation for gyro-averaged distribution function  $\overline{f}$ , take the Vlasov equation and gyro-average  $\rightarrow$  many many pages of maths.

A simple "derivation" is by analogy to the Vlasov equation:

$$\frac{d}{dt}f\left(\mathbf{x},\mathbf{v},t\right)=0$$

Using the chain rule:

$$\frac{\partial f}{\partial t} + \frac{\partial \mathbf{x}}{\partial t} \cdot \frac{\partial f}{\partial \mathbf{x}} + \frac{\partial \mathbf{v}}{\partial t} \cdot \frac{\partial f}{\partial \mathbf{v}} = 0$$
  
and finally putting in  $\frac{\partial \mathbf{x}}{\partial t} = \mathbf{v}$  and the force to get  $\frac{\partial \mathbf{v}}{\partial t}$  gives  
 $\frac{\partial f}{\partial t} + \mathbf{v} \cdot \frac{\partial f}{\partial \mathbf{x}} + \frac{q}{m} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \frac{\partial f}{\partial \mathbf{v}} = 0$ 

# Equation for $\overline{f}$

For the distribution of current loops we now have  $\overline{f}(\overline{\mathbf{x}}, v_{||}, v_{\perp})$ 

• We can choose to use the total kinetic energy K rather than the parallel velocity, and magnetic moment  $\mu$  rather than perpendicular velocity

$$\rightarrow \overline{f}(\overline{\mathbf{x}}, K, \mu)$$

$$K = \frac{1}{2}m\left(v_{||}^2 + v_{\perp}^2\right) \qquad \mu = mv_{\perp}^2/(2B)$$

NB: Not the only possible choice

• Now write down total derivative as before:

$$\frac{d}{dt}f(\mathbf{x},\mathbf{v},t)=0 \qquad \Rightarrow \qquad \frac{d}{dt}\overline{f}(\overline{\mathbf{x}},K,\mu,t)=0$$

 $\sim$ 

From total derivative:

$$rac{d}{dt}\overline{f}\left(\overline{\mathbf{x}},K,\mu,t
ight)=0$$

Expand using chain rule

$$\frac{\partial \overline{f}}{\partial t} + \frac{\partial \overline{\mathbf{x}}}{\partial t} \cdot \frac{\partial \overline{f}}{\partial \overline{\mathbf{x}}} + \frac{\partial K}{\partial t} \frac{\partial f}{\partial K} + \frac{\partial \mu}{\partial t} \frac{\partial f}{\partial \mu} = \mathbf{0}$$

• 
$$\frac{\partial}{\partial t} \overline{\mathbf{x}}$$
 is the motion of the gyro-center  
•  $\frac{\partial}{\partial t} K$  is the change in energy of the particle  
•  $\frac{\partial}{\partial t} \mu \simeq 0$  due to conservation of  $\mu$ 

The motion of the gyro-center is the motion along the magnetic field and the drifts across the magnetic field:

$$\frac{\partial \overline{\mathbf{x}}}{\partial t} = \mathbf{v}_g = \mathbf{v}_{||} \mathbf{b} + \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \frac{1}{\Omega} \left[ \mathbf{v}_{||}^2 \mathbf{b} \times (\mathbf{b} \cdot \nabla) \, \mathbf{b} + \mu \mathbf{b} \times \nabla B \right]$$

The energy of a particle changes due to electric fields:

$$rac{\partial K}{\partial t} = q \mathbf{v}_g \cdot \mathbf{E} + \mu rac{\partial \mathbf{B}}{\partial t}$$

Putting this together gets us...

An equation for particle gyro-centers (current loops)

$$\frac{\partial \overline{f}}{\partial t} + \mathbf{v}_g \cdot \frac{\partial \overline{f}}{\partial \overline{\mathbf{x}}} + \left( q \mathbf{v}_g \cdot \mathbf{E} + \mu \frac{\partial \mathbf{B}}{\partial t} \right) \frac{\partial f}{\partial K} = \mathbf{0}$$

- These move along, and drift (relatively) slowly across, magnetic fields
- The fast gyro-frequency timescale has been removed, so time steps in a simulation can be much larger than for the Vlasov equation
- One velocity dimension removed, reducing the problem to 5D
- But: This is not the gyro-kinetic equation!

- We have neglected the finite size of the Larmor orbits, so assumed that the  $\textbf{E}\times\textbf{B}$  is just given by the E field at the gyro-center position  $\overline{\textbf{x}}$ 
  - $\rightarrow$  Need to average drift around the orbit
- We have not considered how to calculate the E and B fields
   → This is done using Poisson and Ampére laws. Calculation
   of electric field complicated: determined by polarisation, not
   charge separation

There are many subtleties in deriving and using gyro-kinetics, particularly for nonlinear calculations

Whilst the details are more complicated, the principles of gyrokinetic PIC codes are the same as the 1D electrostatic code studied in Comp Lab:

- Gather electrons and ions to calculate gyro-center densities and velocities on grid cells
- Solve for the electric (and magnetic) fields
- Scatter the E and B fields on to the particles. This now involves averaging around a gyro-orbit, typically done by sampling several points on the orbit.
- Galculate the particle drifts, and move the particles
- So to (1)

Many tricks have been developed to reduce the computational cost

#### Simulations of DIII-D using GYRO (left), and of MAST using GS2 (right)



Figures: Waltz et al. Phys. Plasmas (2006), and Hammett PPPL (2002)

#### Results - Dimits shift

- The threshold temperature gradient for significant transport due to ITG turbulence is higher than linear theory predicts
- Non-linear simulations show an initial burst of turbulence, which then dies down to a low level
- Self-regulation of turbulence through generation of mean "zonal" flows



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#### Results - Transport

- For core turbulence, gyro-kinetic codes can now get very close e.g. ITG threshold gradient within 5%
- Fluxes a strong function of gradient, so harder to predict
- Models like TGLF use fits to G-K simulations, and produce quite good results
- A "shortfall" is often observed near the edge, and the cause is still being debated



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## Summary

- There are many instabilities in plasmas which arise because the plasma is composed of different populations of particles, rather than a single homogenous fluid
- These instabilities are mainly electrostatic: magnetic fluctuations are present, and can be an important effect at high β, but are not essential.
- Drift-kinetics and gyro-kinetics average over gyro-motion, removing a fast timescale (cyclotron frequency) and a velocity dimension, making realistic 3D simulations possible

• Still lots of work needed, particularly in extending towards the plasma edge